

The Point Spread Function

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1 Introduction

In this section we discuss the a retinal representation which utilizes geometric, or ray optics to reveal how well parallel light comes to a crisp focus on the retina of the eye. We make use of two standard techniques common for describing the quality of optical systems.

2 Methods

We wish to measure how well the corneas focus light from a single, distant source like a bright star. One standard technique is called the *point spread function* (PSF), which is the retinal energy distribution from a point source [2, 11, 52, 57, 87, 115, 116].

2.1 Point Spread Function

We begin the calculation of the PSF by ray tracing light through the system and onto the retina, as others have done [16, 17, 37, 45, 76, 77, 98, 108]. We assume the point is placed at optical infinity, so that the incoming light rays are parallel. We also assume that the retina is planar, a reasonable approximation given that our entire retinal region of interest (ROI) is 0.64 mm wide and deviates in height from a spherical retinal model by only 3 μm . The PSF is the distribution, or “spread”, of light at the retina. Figure 1 illustrates light rays showering a cornea and forming a PSF on the retina.

We generate a *normalized* PSF (denoted $\overline{\text{PSF}}$), which is simply our PSF divided by the number of rays N_r which land:

$$\overline{\text{PSF}}(x, y) = \frac{\text{PSF}(x, y)}{N_r} \quad (1)$$

and typically plot it as a height field for analysis, as in the example from Figure 2. The $\overline{\text{PSF}}$ for an ideal optical system (like a perfect ellipsoid) would have a very small, localized region of height one and zero elsewhere, as we will see in Figure ???. The *Strehl ratio* is the relative peak intensity of an optical system’s PSF with that of an ideal, aberration-free system. As we are only dealing with geometric optics, we define an equivalent concept, the *geometric* Strehl

Figure 1: Light from a point source is refracted to form the PSF.

ratio, which we will call Strehl_g . Since we normalize our PSF, it is simply the maximum value of our $\overline{\text{PSF}}$.

2.1.1 Sampling

To create the PSF, we must shower the cornea with light rays and trace them all as they are refracted toward the retina. Ideally, every microscopic patch of the cornea would be pierced with a light ray, but for computation purposes, we must sample the continuous surface into discrete corneal points. We use a crosshatched sampling pattern as shown in Figure 3. We typically sample the corneas at a 100×100 resolution (with an inner 99×99 group), which produces 19,801 samples. As most corneal reconstructions are round, approximately 78% or 15,000 of those samples produce valid rays.

We introduce a *simulated pupil* located at the front surface of the eye. As we reduce the diameter of our pupil, we subsequently reduce the number of rays that are allowed to pass. This is consistent with our model that the samples represent a predefined *flux* of incoming light radiation, and culling the peripheral rays with our pupil reduces that flux accordingly. The simple expression for rays allowed to pass is that their distance from the CT axis be less than our pupil radius. That is, if the ray pierces the $z = 0$ plane with polar coordinates (r, θ) , the rays allowed to enter are those satisfying the expression

$$r < r_{pupil}. \quad (2)$$

At each sample in our pupil, we query point and derivative information necessary for determining the refraction into the eye. It is here that we apply our simulated *spectacle correction*. We found that simple aberrations based on astigmatism dominated our distribution, so we compensate by placing a

Figure 2: A height field representation of a sample $\overline{\text{PSF}}$.

Figure 3: The crosshatched sampling pattern used to sample the cornea. Rays from the infinite light source pass through these samples and are refracted into the eye.

simulated infinitely thin contact lens on the cornea to correct for basic cylinder. We search to find the optimal cross-cylinder lens that gives the best retinal focus, which we discuss shortly in Section ??.

The corrected ray then enters the corneal surface according to Snell’s law. After the rays pass through the cornea and eye, they intersect the retinal “plane”, which is a variable distance away. This distance is the third and final parameter in our spectacle correction optimization.

We then sample the retinal plane, dividing it into 257×257 square buckets which are $2.5 \mu\text{m}$, or roughly one-half minute on a side. This produces a square retinal “patch” which is 0.6425 mm on a side. Figure 4 shows a ray striking the retina and landing within one of the sample buckets.

Every time a ray lands in a sample, it contributes one unit to our histogram. The accumulation of all the rays forms the overall PSF energy distribution.

3 See What You See : Simulating Corneal Visual Acuity

4 Introduction

Our goal with this work is to simulate the corneal contribution to visual acuity. In the previous chapter we showed retinal representations of corneal acuity, but these fail to capture what the patient actually sees. We utilize a modified Snellen eye chart and a sample outdoor scene as our input, and image how they

Figure 4: A refracted ray lands in one of the sample buckets on our retinal plane.

would be seen with different patient’s corneas and pupil sizes to achieve a fairly accurate first-person representation of visual acuity.

5 Methods

The use of ray tracing to determine the PSF and the resulting retinal blur for images such as Snellen charts is not new. Section ?? discusses several prominent researchers who have created sophisticated models and optical bench software tools using ray tracing for evaluation of optical performance.

We choose to implement the technique as part of our software suite to provide the final stage of visualization, the simulation of optical acuity through the cornea in question. In this section we’ll discuss the process of calculating the $\overline{\text{PSF}}$, calibrating it with an image, and convolving them together to form the final blurred result.

It is important to note the assumptions and limitations of this technique for the simulation of visual acuity. First, it assumes that all incoming light is parallel, having arrived from optical infinity. Thus, we can only simulate what a patient would see while looking at something reasonably far away. Second, the computational model described in Section ?? is very simplistic and does not take other components of the eye into account, like the lens, vitreous humor or corneal layers. This means the many effects they induce (e.g., ciliary bloom and lens glare) are completely ignored. Third, the pupil only expands to large diameters in extremely low light situations, so simulating the aberrations with an 8 mm pupil on a daylight scene is quite artificial. Finally, as all of our $\overline{\text{PSF}}$ construction uses ray tracing and geometric optics, we ignore the important

Figure 5: Simulation of an eye viewing an image exactly as it was seen by the camera. The centers of projection are aligned and the image is scaled so that one pixel on the image maps to one sample on the $\overline{\text{PSF}}$.

effects of diffraction, which is the limiting acuity factor for small pupils.

5.1 Normalized Point Spread Function

We begin with the normalized point spread function, $\overline{\text{PSF}}$, a computed histogram of retinal energy from a distant point source of light as discussed in Section 2.1. This serves as the “impulse response” of the patient’s optical system. As we mentioned, we sample the retina at half-minute (2.5 micron) intervals; this will be important in the following section on image calibration.

5.2 Image Calibration

When we wish to apply a filter to an image, it is critical that the parameters of the filter be tuned to the spatial frequencies of the image. Our filter is the $\overline{\text{PSF}}$, which is sampled at half-minute intervals, producing a fine grid of retinal energy distribution.

Again, we are given (or synthetically generate) an input image, and our goal is to create a first-person simulation. There are two alternatives, simulating the eye in the camera’s place, or simulating the patient viewing the image from afar. Each of these will affect the calibration differently, and we discuss this in the following sections.

5.2.1 Eye is the Camera

This is perhaps the most effective and convincing simulation, and the one we use to simulate the patient seeing an eye chart in Section 6.1. Here we place the

Figure 6: Simple geometry controls the relationship between an object's distance (d), its image size (h) and the angle it subtends (θ). That relationship is captured by the expression $h = d \tan \theta$.

eye where our camera was when the scene was captured on film as in Figure 5. Before we can accurately simulate the blur the patient would see, we must assure that they are calibrated together, as described below.

A photograph of a scene is taken, or one is synthetically generated. Real or virtual camera information is recorded, e.g., lens, field of view, center of projection, image plane distance, etc. The image is then digitized (or rendered) at the same spatial frequency as our filter.

If we have control over the digitization and camera information, we simply tune our sampling resolution to the correct value. If we do *not* have camera information, but *do* have objects in our scene whose distance to the camera and size we know accurately, we can still adjust our sampling as follows. We want one pixel to be thirty seconds of arc. If an object is at a distance d from our camera, this means (using the simple geometry from Figure 6) that:

$$h_1 = d \tan \theta = d \tan \left(\frac{1}{2} \frac{1}{60} \frac{\pi}{180} \right) \approx d 1.45444 \times 10^{-4} \quad (3)$$

where h_1 represents the distance in object space that we need for one pixel. If an object's size is h , then it should be $\frac{h}{h_1}$ pixels high.

If we do not have either camera or object size and distance information, we have no way to estimate how finely to sample our scene. When we do have this information, but the image has already been digitized, then there are three options. If the image is sampled at calibration density, we're done. If the image is sampled finer than that, then an intelligent bicubic interpolation can be used to reduce the image size without significant loss. However, if the image is sampled more coarsely than calibration, we're in trouble. We either have to reduce the size of our filter or increase the size of our image. Either technique will result in unacceptable artifacts.

If there is no calibrating camera or object information, and the image has already been digitized, we cannot reliably put the eye in the camera. We could, however, simulate what the patient would see if they were to visualize the digitized photo on a computer monitor; this is discussed next.

Figure 7: Simulation of an eye viewing a projected digitized image from afar. The eye’s center of projection is placed at a distance d and the image is scaled so that one pixel on the digital image maps to one sample on the $\overline{\text{PSF}}$.

5.2.2 Patient Views Projected Digital Image

This technique, as shown in Figure 7, allows us to simulate what the patient would see looking at an image from a fixed distance. It is most useful when we do not have object information necessary for eye-is-the-camera calibration.

When images are viewed on monitors in a what-you-see-is-what-you-get (WYSIWYG) fashion, they are represented at 72 dots per inch (DPI). This means one pixel is $\frac{1}{72}$ inch. Plugging in this value for h_1 in Equation 3, we know the distance d must be:

$$d = \frac{h_1}{\tan\left(\frac{1}{2} \frac{1}{60} \frac{\pi}{180}\right)} \approx 95.5 \text{ inches} \approx 8 \text{ feet.} \quad (4)$$

Therefore, if we perform our filtering on an unmodified input image, the result would be what a person would see viewing the picture on a monitor from a distance of approximately 8 feet. This is the technique we employ with our outdoor scene from Section 6.2.

5.3 Measuring Visual Acuity

Visual acuity is the measurement of the eye’s ability to resolve the form and detail of an object. It is most commonly determined by testing the ability of the patient to read standard letters at a fixed distance [10, 25].

5.3.1 Snellen Notation

In 1862 Snellen devised a system for measuring acuity which has since become a fundamental clinical technique for acuity assessment [25]. The smallest detail of an object an eye can see is called the *minimum angle of resolution* (MAR) of the eye. The *Snellen fraction* is the reciprocal of the MAR. The test is usually performed at 20 ft (6 m), an acceptable approximation to optical infinity.

Figure 8: A typical Snellen letter.

<i>MAR</i>	<i>Snellen decimal notation</i>	<i>Snellen fraction</i>
0.5	2	20/10
1	1	20/20
2	1/2	20/40
4	1/4	20/80
8	1/8	20/160
⋮	⋮	⋮

Table 1: The relationship between MAR and Snellen notations.

On a typical chart, letters are scaled so that each will subtend an angle of 5 minutes at a given distance. The details of the letters themselves make an angle of 1 minute of arc, as shown in the classic “E” in Figure 8. The letters are then labeled by this distance, e.g., the “20-foot” letter makes a total angle of 5 minutes at a distance of 20 feet. The classic fraction is recorded as the ratio of the testing distance and the label of the smallest letter a patient can resolve. Thus, if a patient at 20 feet is only able to read the “100 foot” letter, their vision would be classified as “20/100”. Table 1 shows the relationship between MAR and Snellen notation. A historical note: one minute of arc became a standard in the days of early astronomers who determined it to be the minimum angular separation for two different stars to be perceived as distinct [25].

5.3.2 Simulated Snellen Eye Chart

Tumbling E charts like the one shown in Figure 9 are often used for preliterate children. We choose to use a modified version of this instead of a standard Snellen chart since rendering the letter “E” to an image requires no anti-aliasing as would letters with curved edges. We are able to create Es down to 5×5 pixels with no loss.

Our test acuity image is shown in Figure 10, which has been calibrated so that each pixel is thirty seconds. This is the spacing between the bars of the smallest E, used for testing 20/10 acuity. We include Es which double in size

Figure 9: A typical acuity chart with tumbling Es used for acuity assessment of preliterate children. This chart is especially useful for us since rasterization requires no anti-aliasing as would a chart with curved-edge letters.

Figure 10: The image we use for our acuity simulation. The size of the Es ranges from 20/10 on the left to 20/160 on the right. The fan pattern on the far right is used to test for astigmatism.

Figure 11: The image we use for our outdoor scene simulation, a photograph of U. C. Berkeley's Campanile courtesy of Paul Debevec.

up to 20/160. We add a fan pattern to test for astigmatism, whose direction is determined by the bars perceived as least blurred.

5.3.3 Outdoor Scene

As shown in Figure 11, the input image is a crisp low dynamic range digital photograph of U. C. Berkeley's Campanile tower, courtesy of Paul Debevec. It has clearly defined edges and a full-range luminance histogram. The simulations we present are what a patient would see when looking at the full 250×500 pixel image on a 72-dpi computer screen at a distance of 8 feet. The blur is identically computed for each of the R, G and B channels in the image, so there is no chromatic aberration in the simulation. There was no digital processing done on the input image; every pixel was simply copied from the original PhotoCD, down to the small smudge on the right of the tower.

Figure 12: Instead of convolving the $\overline{\text{PSF}}$ with our image (shown in grey), we multiply their Fourier transforms and return the inverse Fourier transformation (shown in black).

5.4 Convolution

Now that we have our input images and $\overline{\text{PSF}}$ “impulse response” distribution, we simply need to convolve them together to form the blurred output images. We make use of the *convolution theorem* which tells us the convolution in the spatial domain can be obtained by taking the inverse Fourier transform of the products of the spectra in the frequency domain, as shown in Figure 12. That is,

$$\begin{aligned} \text{Image}_{\text{blur}}(x, y) &= \overline{\text{PSF}}(x, y) * \text{Image}(x, y) \\ &= \mathcal{F}^{-1} \{ \mathcal{F} \{ \overline{\text{PSF}}(x, y) \} \times \mathcal{F} \{ \text{Image}(x, y) \} \} \end{aligned} \quad (5)$$

where \mathcal{F} is the Fourier transform and \mathcal{F}^{-1} is the inverse Fourier transform [34].

6 Results

We compile the results of simulating all of the corneas with pupil sizes of 2, 4 and 8 mm viewing both our test images in Section ??, and discuss the overall results in the following sections.

6.1 Snellen Eye Chart

As anticipated, the corneas have much better vision with small pupils than with large. The regular corneas (those without PRK, keratoconus or monocular diplopia) have excellent spectacle-corrected vision with no astigmatism even up to 8 mm; acuity is estimated to be between 20/10 and 20/40. The problem eyes have acute loss of contrast and acuity, sometimes even with small pupils. A telltale ghost image forms with our 8 mm monocular diplopia eye, situated

about $\frac{4}{3}^\circ$ above the primary image. Overall acuity ranges at around 20/80 for the PRK and diplopic eye to worse than 20/160 for our keratoconic eye.

6.2 Outdoor Scene

The aberrations that we witnessed with the simulated Snellen chart were more mild than with our outdoor scene, since our scene did not have a comparable degree of contrast and sharp edges, and thus was more forgiving. In general, it was harder to differentiate the blur from different corneas, as the results all seemed to converge with large pupils. The most striking feature was the “muddying” of the scene, as everything in the interior of the tower blurred to a dark grey. The ghost image was not as distinct for the diplopic eye as it had been for our Snellen chart.

7 Conclusion

We presented a technique for the simulation of first-order visual acuity using a precomputed normalized point spread function of the eye. We utilized two sample input images: a modified tumbling E Snellen chart with an astigmatic test fan, and a sample full-color outdoor scene. Our results showed a fair approximation of visual acuity, with expected increased blur and loss of contrast for larger pupils as peripheral aberrations became more dominant.

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